Energy and water vapor transport across a simplified cloud-clear air interface

L. Gallana, S. Di Savino, F. De Santi, M. Iovieno, D. Tordella

Dipartimento di Ingegneria Meccanica e Aerospaziale, Politecnico di Torino, Corso Duca degli Abruzzi 24, 10129 Torino, Italy

E-mail: luca.gallana@polito.it, daniela.tordella@polito.it

Abstract. We consider a simplified physics of the could interface where condensation, evaporation and radiation are neglected and momentum, thermal energy and water vapor transport is represented in terms of the Boussinesq model coupled to a passive scalar transport equation for the vapor. The interface is modeled as a layer separating two isotropic turbulent regions with different kinetic energy and vapor concentration. In particular, we focus on the small scale part of the inertial range as well as on the dissipative range of scales which are important to the micro-physics of warm clouds. We have numerically investigated stably stratified interfaces by locally perturbing at an initial instant the standard temperature lapse rate at the cloud interface and then observing the temporal evolution of the system. When the buoyancy term becomes of the same order of the inertial one, we observe a spatial redistribution of the kinetic energy which produce a concomitant pit of kinetic energy within the mixing layer. In this situation, the mixing layer contains two interfacial regions with opposite kinetic energy gradient, which in turn produces two intermittent sublayers in the velocity fluctuations field. This changes the structure of the field with respect to the corresponding non-stratified shearless mixing: the communication between the two turbulent region is weak, and the growth of the mixing layer stops. These results are discussed and compared with laboratory and numerical results with and without stratification.

1. Introduction

Warm clouds as stratocumuli swathe a significant part of earth's surface and play a major role in the global dynamics of atmosphere by strongly reflecting incoming solar radiation – thus contributing to the Earth's albedo – so that an accurate representation of their dynamics is important in large-scale analyses of atmoshperic flows [1]. They are controlled by the tight interplay between radiative driving, turbulence, surface fluxes, latent heat release, and entrainment. Among them, the mixing and entrainment processes at the cloud top have been identified as fundamental to determine the internal structure of warm clouds, so that a clear and complete understanding of their physics is required [2]. As pointed out by Malinowsky et al. [3], data from most field campaigns and large-eddy simulations are too poorly resolved to allow to infer the details of the interfacial layer, even if they indicate that, in order to allow for entrainment, a high level of turbulence must be present. For this reason, in this work we study the local transport through a clear air/cloud interface through DNS (Direct Numerical Simulation). As our focus is on the dynamics of the smallest scales of the flow which influence the microphysics of warm clouds, we have simulated an idealized configuration to understand, under controlled conditions, some of the basic phenomena which occur at the cloud interface

Figure 1. Kinetic energy spectra from in situ atmospheic measurements [7, 8, 9, 10] compared with present work spectrum.

over length scales of the order of few meters. In these conditions, we solve scales from few meters down to few millimeters, that is, we resolve only the small scale part of the inertial range and the dissipative range of the power spectrum in a small portion $(6 \text{ m} \times 6 \text{ m} \times 12 \text{ m})$ of the atmosphere across the clou - clear air interface. This allows us to investigate the dynamics of entrainment which occurs in a thin layer at the cloud top, which a smaller scale with respect ot the scale explicitly resolved in large-eddy simulations of clouds [4]. In this preliminary work, we focus on two concomitant aspects of the cloud top mixing layer: the effect of the presence of a stratification and of a turbulent kinetic energy gradient. We do not consider the wind shear neither the phenomena linked to the processes of evaporation and condensation and radiative cooling which are important in conditions of buoyancy reversal [5, 6]. Therefore, our simulations were performed by applying the Boussinesq approximation to the Navier-Stokes momentum and energy equations together with a passive scalar transport equation which models the water vapor transport.

The details of the physical problem we have considered and of the governing equations are given in section 2. Section 3 contains a selection of our main results about intermittency, energy redistribution and entrainment. Conclusion remarks are in Section 4.

2. The physical problem

We consider the interface between clear air and cloud in a 6 $m \times 6$ $m \times 12$ m parallelepiped domain. In this way we are able to simulate, compared to in situ measurement of atmospheric energy spectra, shown in figure 2, the lowest part of the inertial range and the dissipative one. The system is composed, as shown in figure 2, by two homogeneous and isotropic turbulent regions that interact through a mixing layer, whose initial thickness has been set of the same order of the integral scale of the turbulence background ℓ , here assumed equal to $3 \cdot 10^{-2}$ m. The two isotropic regions have a different kinetic energy: within the cloud the kinetic energy is considered to be higher with respect to the external region. The root mean square of the velocity inside the cloud is $u_{rms} = 0.2$ m/s, and the energy ratio with the external region is $E_{int}/E_{ext} = 6.7$. This energy ratio is of the same order of the ones measured in warm clouds (see, e.g., [3]) and, furthermore, allow us to compare with laboratory and numerical experiments on shearless mixing (see [11, 12]) in absence of any stratification. In our simulations the Taylor microscale Reynolds number Re_λ of the higher energy region is equal to 250. Buoyancy is taken into account through a local perturbation in the profile of density distribution inside the troposphere, and the water vapor is considered as a passive scalar. The Prandtl $(Pr= 0.74)$ and the Schmidt number ($Sc = 0.61$) considered refers to an altitude of 1000 m s.l.

Figure 2. Scheme of the initial conditions. E_1 is the mean initial turbulent kinetic energy of the bottom region (inside the cloud), E_2 of the upper region (outside the cloud). For top cloud mixing $E_1/E_2 = 6.7$, while for bottom mixing $E_1/E_2 = 6.7$. The water vapor is initially present in the bottom region. The zoom in the yellow circle is an example of initial temperature perturbation of the standard boundary layer lapse rate.

Table 1. Initial stratification level parameters. G is the maximum gradient of θ' , expressed in terms of the standard troposphere gradient $G_0 = 0.0065$ Km^{-1} ; $N_{ci} = \sqrt{\alpha T_0 g \frac{dT}{dx}}$ $\frac{dT}{dx_3}$ is the characteristic Brunt-Väisälä frequency of initial condition. The Froude number Fr= $\frac{u'_{rms}}{N_{ic}\ell}$ and the Reynolds Buoyancy number Re_b = $\frac{\varepsilon N^2}{\nu}$ $\frac{N^2}{\nu}$ give a measure of the order of magnitude of the buoyancy forces compared with the inertial terms.

G	ΔT [K]	N_{ic} [s ⁻¹]	Fr	Rеь
$2G_0$	$4.0e-3$	$2.13e-2$	31.2	
30G ₀	$6.0e-2$	$5.24e-2$	12.7	112
100G ₀	$2.0e-1$	$1.50e-1$	4.4	273
500G ₀	1.0e0	$3.35e-1$	1.8	833
$5000G_0$	1.0e1	1.06e0	0.62	2635

The initial conditions for the temperature perturbation can be found in table 1. The ratio between inertial and buoyancy forces is expressed by the Froude number Fr , based on the maximum gradient within the initial interface, which ranges from 31.2 (negligible stratification) to 0.62 (strong stratification).

We use the continuity, momentum and energy balance equations within the Boussinesq approximation which considers the buoyancy effects for small temperature variations [13]:

Figure 3. Spatial distribution along the vertical coordinate x_3 of the kinetic energy (a), temperature variance (b) and water vapor concentration variance (c) at $t/\tau \approx 6$ for different levels of stratification.

$$
\nabla \cdot \mathbf{u}' = 0 \tag{1}
$$

$$
\frac{\partial \mathbf{u}'}{\partial t} + (\mathbf{u}' \cdot \nabla) \mathbf{u}' = -\nabla \frac{\tilde{p}}{\rho} + \nu \nabla^2 \mathbf{u}' + \alpha \mathbf{g} \theta'
$$
\n(2)

$$
\frac{\partial \theta'}{\partial t} + \mathbf{u'} \cdot \nabla \theta' + u_3 G = \kappa \nabla^2 \theta'
$$
\n(3)

$$
\frac{\partial \chi}{\partial t} + \mathbf{u}' \cdot \nabla \chi = \kappa \nabla^2 \chi,\tag{4}
$$

where the temperature θ is composed as the sum of a fluctuation $\theta'(\mathbf{x},t)$, a static component $\tilde{\theta}(x_3) = Gx_3$, and a reference constant temperature θ_0 , $\tilde{p} = p + \alpha g x_3 (T_0 + Gx_3/2)$, u' is the velocity fluctuation and χ is the vapor concentration of the air - water vapor mixture present in the cloud, here considered as a passive scalar. These equations are solved using a Fourier-Galerkin pseudospectral method with a fourth order Runge-Kutta time integration [14].

3. Results

According to the ratio between buoyancy force and kinematic forces (that are advection and diffusion), the evolution of the system can be split in two main stages. As long as the ratio remains small, there are no significant differences with respect to a non-stratified case. On

Figure 4. Spatial distribution in the vertical direction x_3 of the skewness and the kurtosis of vertical component of the velocity (a - b), temperature (c - d) and water vapour concentration (e - f) at $t/\tau = 6$ for different levels of stratification.

contrary, as the stratification perturbation level become higher, buoyancy effects are no more negligible: differences are present from both a quantitative and qualitative point of view. These consideration can be observed in figure $3(a - b - c)$, where the second order moment of velocity, temperature, and vapor are shown. For low stratification levels there are no significant differences, while for intensive stratification differences are visible. In particular, in correspondence of the local temperature perturbation, the formation of a layer with a pit of kinetic energy can be observed. The presence of such a layer deeply changes the physics of the system, because in this situation two interfaces are produced. The first interface, (which is present also in the absence of stratification), separates the high turbulent energy region from the pit, while the second one (not present without stratification) separates the low turbulent energy region from the center of the mixing layer. Therefore, the stratification induces a physical separation between the two external regions, decreasing greatly the interaction between them. Both interfaces presents an intermittent behavior, as shown with normalized third and fourth moments profiles in figures $4(a - b)$. In the highly stratified case, in fact, two peaks of skewness and kurtosis can be observed: one is placed inside the cloud, and the other is nearly placed where the intermittency peaks without stratification, see $Fr = 0.6$ in figure 4 (a). Observing the magnitude of the kurtosis maximum in figure 4 (b), it can be noted that the peak inside the cloud reaches values as high as 4, that are about the 10% larger than when the stratification is milder.

The behavior of the temperature and the water vapor fields appear qualitatively analogue as far as the shape of the statistical distributions are concerned, but are quantitatively different. See, for instance, the temperature and the vapor concentration peaks in the skewness and kurtosis distribution, shown in 4 (c -d -e- f).

It can be also observed that higher levels of stratification produce a relevant reduction of intermittency in the flow, with a reduction of about 70% in the peaks of skewness and kurtosis. The interaction between the two regions aside the interface is greatly reduced, so the fluctuation at the sides of the mixing layer are damped, preventing the formation of the intermittent layer typical of the passive scalar transport [15, 16]. Moreover, observing both figures 3 and 4, it is clear that the thickness of the mixing layer is reduced in case of high stratification (see next section for more details).

3.1. The onset of a kinetic energy pit

As shown in figure 5 (a), in case of high stratification level in the center of the domain, in correspondence with the temperature step, the birth of a layer with very low levels of kinetic energy can be observed, obtaining a pit of kinetic energy. Varying the stratification level, the genesis and the evolution of such pit can be measured by considering the ratio ζ between the variance of vertical velocity in the low energy region (mean value) and in the center of the pit (where the variance reaches its minimum), as shown in figure 5 (b). Qualitatively similar results have been observed in the large-eddy simulations of stratocumulos-topped planetary boundary layer carried out by different physics of the atmosphere research group, as described in [17, fig.7 at pag. 11].

A visualization of such phenomenology is represented in figure 6, where the vertical velocity fluctuations in a vertical slice of the domain are represented using an elevation plot (where such elevation is proportional to the square of u₃). In presence of a mild stratification, $Fr = 12$, even after 8 time scale, there is a smooth mixing layer between the high (left) and the low (rigth) energy regions. The differences in the case of strong stratification, with $Fr = 0.62$, after 4 time-scales (b) and 8 time scales (c), are clearly visible: a separation layer is present in the center of the domain, that becomes even more evident as the time pass by.

The presence of the pit generates a physical separation between the two external regions, by damping the turbulent mixing, and thus reducing the exchange of information. As a consequence,

Figure 5. (a) Temporal evolution of the spatial distribution along the vertical direction x_3 of the turbulent kinetic energy when Fr=1.8. A pit of energy appears after about 6 eddy turnover times in correspondence of the original interface.

(b) Temporal evolution of the ratio ζ between the mean vertical velocity variance in the lower energy (clear air side) region u_{3MLE} and the minimum value of the vertical velocity u_{3MIN} . When this ratio departs from 1, a pit of kinetic velocity appears as shown in (a).

Figure 6. Visualization of the vertical component of the velocity in a vertical plane. The elevation is proportional to the square of vertical velocity fluctuations and the colors respect the velocity directions (blue for downward , red for upward) in various case: $Fr = 12$ after 8 eddy turnover times (a) and for Fr=0.62 after 4 and 8 eddy turnover times (panels b, c respectively). In the last two panels it can be observed the formation of the pit of kinetic energy.

Figure 7. Thickness of the mixing layer of velocity (a) and temperature (b). For both fields, the mixing layer thickness is the same as defined in [16].

there is a saturation of the thickening of the mixing layer Δ_E . As shown in figure 7 (a), there is an interruption of the growth of Δ_E .

Looking to the temperature mixing layer thickness Δ_{θ} , shown in figure 7 (b), it can be seen that, for strong stratification, the thickening stops approximately after the same amount of times scale required by Δ_E . In that case, contrary to what observed for the kinetic energy, the thickening does not stop suddenly, but rather with a transient that lasts a couple of time scales

3.2. Entrainment

The entrainment of clear air inside the cloud is an important aspect of the top cloud interface as it concurs in the evaporation/condensation of droplets inside a cloud [1].

In any plane parallel to the interface, in absence of a mean velocity, only downward velocity fluctuations can transport clear air into the cloud. Their presence can represented by a marker function ψ that is equal to 1 where u_3 is negative, and 0 otherwise. Its average in each horizontal plane, shown in figure 8 (a), shows a small deviation from the mean value of 0.5 which would be observed in an homogeneous and isotropic flow. This implies that upward and downward fluctuations are almost equally distributed; the maximum departure from a symmetric distribution is about 4%, and the spatial distribution of ψ looks like the one which have been observed in the third order moment of the velocity, see figure 4 (a).

The entrainment of clear air is responsible of the growth of the cloud. In fact, the velocity $w_e = dx_{3,i}/dt$ of the cloud top interface $($x_{3,i}>$ is the mean vertical position of the cloud top,$ obtained as the position where the mean vapor concentration is 25%) has often been used as a parameter to measure the entrainment rate [5, 4]. In figure 8 (b), the time decay of w_e for different levels of stratification perturbation. In presence of weak stratification, that is Fr larger than 4, its value gradually decreases with an almost exponential law, due to the decay of the kinetic energy; on the contrary, when the stratification perturbation is stronger, the damping of w_e is much faster, and the entrainment vanishes after few times scale, when the presence of the pit of kinetic energy substantially reduces the flux of clear air inside the cloud.

Figure 8 (c) shows the vertical derivative of the downward flux of clear air when $Fr = 1.8$ The downward flux reduces as the flow evolves and its derivative, which represents the net variation of $1 - \chi$ at a given instant, rapidly tends to zero inside the cloud; this implies that the entrainment of clear air is confined to a thin interfacial layer.

Figure 8. Evolution of the entrainment across the top cloud. Fraction ψ of downwards velocity (a) and vertical variation of the mean flux of clear air into the cloud (c) when with $Fr = 1.8$ Evolution of entrainment velocity w_e normalized with the high kinetic energy E_1 root mean square (b), see figure 2.

4. Conclusion

In this work we have carried out numerical simulations of the transport of energy and scalars in a turbulent shearless mixing layer associated to temporal perturbation of the temperature lapse rate across the clear air - cloud interface. The perturbation introduce locally a stable stratification. This idealized configuration models some of the phenomena which are present in the kinetic dinamics of the cloud and clear air interaction, namely those linked to turbulent mixing and entrainment.

We have shown that this flow configuration develops an horizontal layered structure characterized by a low kinetic energy sublayer which acts as a barrier for the transport between the cloud and the external ambient. In our transient simulations, this flow structure appears when the buoyancy terms becomes of the same order of magnitude of the inertial ones, therefore the time needed for this transition becomes shorter when the stratification is more intense. Once buoyancy dominates, there is a spatial redistribution of the kinetic energy which produces a central sublayer where the kinetic energy is very low. In this condition, we observed two highly intermittent regions with opposite kinetic energy gradients. As a direct consequence, the entrainment is damped.

Bibliography

- [1] Wood R 2012 Mon. Weather Rev. 140 2373–2423 ISSN 0027-0644
- [2] Gerber H, Frick G, Malinowski S P, Jonsson H, Khelif D and Krueger S K 2013 J. Geohys. Res. D 118 12094–12109 ISSN 2169-897X
- [3] Malinowski S P, Gerber H, Jen-La Plante I, Kopec M K, Kumala W, Nurowska K, Chuang P Y, Khelif D and Haman K E 2013 Atmos. Chem. Phys. 13 12171–12186 ISSN 1680-7316
- [4] Moeng C 2000 J. Atmos. Sci. 57 3627–3643 ISSN 0022-4928
- [5] Mellado J P 2010 J. Fluid Mech. 660 5–36 ISSN 0022-1120
- [6] Mellado J P, Stevens B and Schmidt H 2014 J. Atmos. Sci. 71 1040–1057 ISSN 0022-4928
- [7] Biona C, Druilhet A, Benech B and Lyra R 2001 Agric. Forest Meteor. 109 135–141 ISSN 0168-1923
- [8] Katul G, Geron C, Hsieh C, Vidakovic B and Guenther A 1998 J. Appl. Meteor. 37 1533–1546 ISSN 0894-8763
- [9] Lothon M, Lenschow D H and Mayor S D 2009 Bound. Lay. Meteor. 132 205–226 ISSN 0006-8314
- [10] Radkevich A, Lovejoy S, Strawbridge K B, Schertzer D and Lilley M 2008 Quart. J. Roy. Meteor. Soc. 134 317–335 ISSN 0035-9009
- [11] Veeravalli S and Warhaft Z 1989 J. Fluid Mech. 207 191–229
- [12] Tordella D and Iovieno M 2011 Phys. Rev. Lett. **107** 194501
- [13] Drazin P and Reid D 1981 Hydrodynamic stability (Cambridge (UK): Cambridge University Press)
- [14] Iovieno M, Cavazzoni C and Tordella D 2001 Comp. Phys. Comm. 141 365–374
- [15] Ma B and Warhaft Z 1986 Phys. Fluids 29(10) 3114–3120
- [16] Iovieno M, Di Savino S, Gallana L and Tordella D 2014 J. Turb. 15 311–334 ISSN 1468-5248
- [17] Moeng C, Cotton W, Bretherton C, Chlond A, Khairoutdinov M, Krueger S, Lewellen W, MacVean M, Pasquier J, Rand H, Siebesma A, Stevens B and Sykes R 1996 Bull. Am. Meteorol. Soc. 77 261–278 ISSN 0003-0007